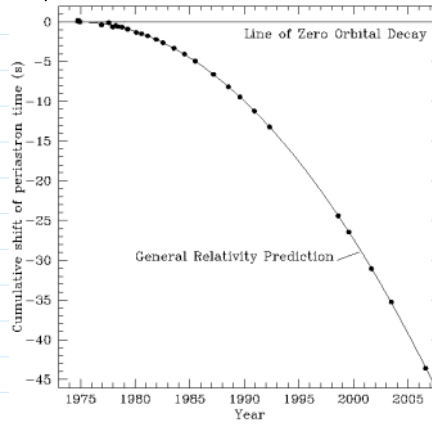
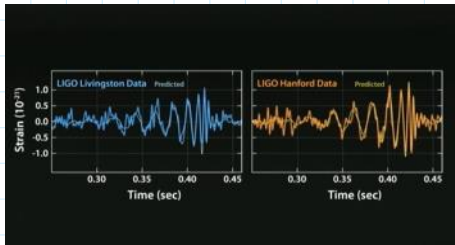


Relativistic Cosmology

G.R.: - predicted the exact perihelion shift of Mercury (43"/century of the 574"/century)
 - predicted light will be deflected by gravity



- gravitational waves



GW091415

Differences b/w Newtonian Grav. & GR

- 1) GR is based on 10 potentials, not 1
- 2) GR is a non-linear theory, i.e. total effect of several bodies is not the simple sum of separate effects
- 3) Pressure is 'natively' a source of gravitation as well as density
- 4) GR usually expressed in terms of geometry;

the 10 potentials give the metric of space-time

The Way of Newton:

Mass tell gravity how to exert a force
($F = -GMm/r^2$)

Force tells mass how to accelerate ($F=ma$)

The Way of Einstein:

Mass-energy tells space-time how to curve

Curved space-time tells mass-energy how to move.

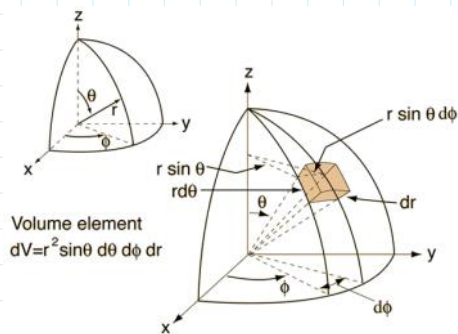
Spacetime characterized by metric

In SR, let x, y, z denote the Cartesian coordinate system and t is the time:

Let 2 neighbouring points in space-time be x, y, z, t and $x+dx, y+dy, z+dz, t+dt$

$$ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2 = -c^2 dt^2 + dl^2$$

If $ds=0$, then $\frac{dl}{dt} = c$; ie, set $ds=0$ to describe behavior of light



In spherical polar coordinates,

$$x = r \sin \theta \cos \phi$$

$$y = r \sin \theta \sin \phi$$

$$z = r \cos \theta$$

$$dl^2 = dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$$

is SR, $ds^2 = -c^2 dt^2 + dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$

flat, Euclidean space, Minkowski metric

flat, Euclidean space, Minkowski metric

in GR, $ds^2 = \sum_{\mu, \nu} g_{\mu, \nu} dx^\mu dx^\nu$ in general

where $g_{\mu\nu} = \begin{pmatrix} - & - & - & - \\ & - & - & - \\ & & - & - \\ & & & - \end{pmatrix}$ is the metric w/ 10 free parameters/potentials

$\mu, \nu = 0, 1, 2, 3$; x^0 is the time coord.

and x^1, x^2, x^3 are the spatial coord's.

We can simplify by imposing the cosmo. principle. So, can remove the x -terms and spatial part has a constant curvature

$$dl^2 = \frac{dr^2}{1-Kr^2} + r^2(d\theta^2 + \sin^2\theta d\phi^2)$$

Now, add time-dependence to dl to allow space to grow or shrink:

Robertson-Walker Metric

$$ds^2 = -c^2 dt^2 + a^2(t) \left[\frac{dr^2}{1-Kr^2} + r^2(d\theta^2 + \sin^2\theta d\phi^2) \right]$$

	SR	GR/RW
Element of radial distance	dr	$\frac{adr}{\sqrt{1-Kr^2}}$
Element of circumference ($\sin\theta = 1$)	$r d\phi$	$a r d\phi$
Element of area dA	$r^2 \sin\theta d\theta d\phi$	$a^2 r^2 \sin\theta d\theta d\phi$
Element of volume dV	$r^2 \sin\theta d\theta d\phi dr$	$\frac{a^3 r^2 \sin\theta d\theta d\phi dr}{\sqrt{1-Kr^2}}$

here, r, θ, ϕ are comoving coords.

The Proper Distance

The distance to a galaxy far away at $r=r_1$, $\theta=\theta_1$, $\phi=\phi_1$ at time t_1 ,

$$dl = \frac{a dr}{\sqrt{1-kr^2}} \rightarrow l_1 = a(t_1) \int_0^{r_1} \frac{dr}{\sqrt{1-kr^2}} = a(t_1) f(r_1)$$

Compute the same thing at time t_2

$$l_2 = a(t_2) f(r_1)$$

$$\text{i.e. } \frac{l_2}{l_1} = \frac{a(t_2)}{a(t_1)}$$

- meaning of $a(t)$ in GR is as a scale factor (same as Newt. case)
- Since r_1 is a comoving coord:
$$\dot{l}_1 = \dot{a}(t_1) f(r_1) = \dot{a}(t_1) \frac{l_1}{a(t_1)} = H(t_1) l_1$$

or, at the current time $v_p(t_0) = H_0 l(t_0)$



