

Theory Notes for AGNs

1) The inner radius of accretion disks is limited to the innermost stable orbits which are derived from G.R., depend on spin (a)

Using $G=c=1$, the natural unit of length used in AGN studies

$$\text{is } r_g = \frac{GM}{c^2} = M \quad (0M \leq a \leq 0.998M)$$

The innermost bound circular orbits are

$$r_{ms} = M \left\{ 3 + Z_2 \mp \left[(3 - Z_1)(3 + Z_1 + 2Z_2) \right]^{1/2} \right\}$$

where upper sign is prograde orbits

lower sign is retrograde orbits

$$Z_1 = 1 + \left(1 - \frac{a^2}{M^2} \right)^{1/3} \left[\left(1 + \frac{a}{M} \right)^{1/3} + \left(1 - \frac{a}{M} \right)^{1/3} \right]$$

$$Z_2 = 3 \left(\frac{a^2}{M^2} + Z_1^2 \right)^{1/2}$$

$$\text{For } a=0 \text{ (Schwarzschild)} \quad r_{ms} = 6M = 6 \frac{GM}{c^2} = 3R_{sch.}$$

$$a=M \text{ (max. Kerr BH)}, \quad r_{ms} = \begin{cases} M & \text{prograde} \\ 9M & \text{retrograde} \end{cases}$$

The binding energy of the marginally stable circular orbit increases as $a \rightarrow M$. So, accretion efficiency $\eta = 0.057 \rightarrow 0.42$

2) Effects of GR on α -disk theory

Recall, we had the dissipation per unit disk face area

$$D(R) = \frac{3GM\dot{M}}{8\pi R^3} \left(1 - \left(\frac{R_*}{R} \right)^{1/2} \right)$$

↳ changed this term to absorb the GR corrections

$$\text{now, } \Omega(R) = \frac{3GMM}{8\pi c R^3} R_E(x) \text{ where } x = \frac{R}{r_g}$$

also, $\sqrt{\Sigma} = \frac{M}{3\gamma} \left(1 - \left(\frac{R_E}{R}\right)^{1/2}\right)$ from cons. of ang-mom.

$$\text{becomes } \sqrt{\Sigma} = \frac{M}{3\gamma} R_T(x)$$

Hydrostatic equilibrium in the vertical direction is

$$\frac{dP}{dz} = \rho g_z = - \frac{GM_p z}{R^3} R_z(x)$$

$$\text{where: } R_E(x) = \frac{C(x)}{B(x)}, \quad R_T(x) = \frac{C(x)}{A(x)}, \quad R_z(x) = x^{-1} \left[L^2 - a_*^2 (E_\infty - i) \right]$$

$$A(x) = 1 - \frac{2}{x} + \frac{a_*^2}{x^2}, \quad B(x) = 1 - \frac{3}{x} + \frac{2a_*}{x^{3/2}} \quad \text{and}$$

$$\begin{aligned} C(x) = & 1 - \frac{Y_{ms}}{Y} - \frac{3a_*}{2Y} \ln\left(\frac{Y}{Y_{ms}}\right) - \frac{3(Y_1 - a_*)^2}{YY_1(Y_1 - Y_2)(Y_1 - Y_3)} \ln\left(\frac{Y - Y_1}{Y_{ms} - Y_1}\right) \\ & - \frac{3(Y_2 - a_*)^2}{YY_2(Y_2 - Y_1)(Y_2 - Y_3)} \ln\left(\frac{Y - Y_2}{Y_{ms} - Y_2}\right) - \frac{3(Y_3 - a_*)^2}{YY_3(Y_3 - Y_1)(Y_3 - Y_2)} \ln\left(\frac{Y - Y_3}{Y_{ms} - Y_3}\right) \end{aligned}$$

where $Y = \sqrt{x}$; Y_{ms} = value of y at r_{ms}

$$Y_{1,2,3} = 3 \text{ roots of } Y^3 - 3Y + 2a_* = 0$$

$$L(x) = \frac{x^{1/2} (1 - 2a_* x^{3/2} + a_*^2 x^{-2})}{B''(x)}, \quad E_\infty(x) = \frac{1 - 2/x + a_* x^{-3/2}}{B''(x)}$$