

$$\text{We had } \frac{dW}{dAd\omega} = \frac{1}{c^3} \frac{\omega^4}{R^2} |\hat{d}(\omega)|^2 \sin^2\theta$$

To get the total energy per freq. interval, integrate over the surface of sphere  $R$

$$\frac{dW}{d\omega} = \frac{8\pi\omega^4}{3c^3} |\hat{d}(\omega)|^2 \quad \leftarrow \text{Only for non-rel. particles}$$

$\therefore$  The spectrum of dipole radiation is  $\propto$  the freq. of the oscillation of the dipole moment.

### Relativistic Larmor Formula

The emitted power is invariant under Lorentz transformations:  $P = \frac{dW}{dt}$  (lab frame) =  $\frac{dW'}{dt'}$  (rest frame)

(rest frame)

albeit momentarily and moving non-relativistically

$$\text{So, } P = P' = \frac{2q^2}{3c^3} |\vec{a}'|^2$$

It can be shown that  $a'_{\parallel} = \gamma^3 a_{\parallel}$  and  $a'_{\perp} = \gamma^2 a_{\perp}$   
 $(\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}})$

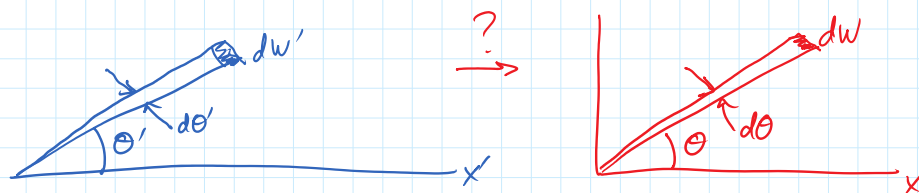
$$P = \frac{2q^2 \gamma^4}{3c^3} (a_{\perp}^2 + \gamma^2 a_{\parallel}^2)$$

A small acc'n of a rel. particle leads to a large energy loss!

### Angular Dist'n of Emitted & Received Power

Consider an amount of energy  $dW'$  that is emitted into the solid angle  $d\Omega' = \sin\theta' d\theta' d\phi'$  about the

direction at angle  $\theta'$  to the  $x'$ -axis



$$\text{let } \mu = \cos\theta \quad ; \quad \mu' = \cos\theta'$$

$$\text{so that } d\Omega = d\mu d\phi \quad ; \quad d\Omega' = d\mu' d\phi'$$

Recall that energy & momentum form a four-vector  
so the transformation of the energy of the radiation is

$$dW = \gamma(dW' + v dp'_x) = \gamma\left(1 + v \frac{dp'_x}{dW'}\right) dW' = \gamma\left(1 + \frac{v c \mu'}{c dW'}\right) dW'$$

$$dW' = dp'c \rightarrow dp'_x = \cos\theta' dp' = \cos\theta' \frac{dW'}{c} = \mu' \frac{dW'}{c}$$

$$dW = \gamma(1 + \beta \mu') dW' \quad \left(\beta = \frac{v}{c}\right)$$

From the aberration of light,

$$\mu = \frac{\mu' + \beta}{1 + \beta \mu'}$$

$$\begin{aligned} \text{differentiating, } \frac{d\mu}{d\mu'} &= (1 + \beta \mu')^{-1} + (\mu' + \beta)(-1)(1 + \beta \mu')^{-2} \beta \\ &= \frac{1 + \beta \mu' - \beta \mu' - \beta^2}{(1 + \beta \mu')^2} = \frac{1 - \beta^2}{(1 + \beta \mu')^2} \end{aligned}$$

$$\text{but } \gamma = \frac{1}{\sqrt{1 - \beta^2}} \rightarrow 1 - \beta^2 = \frac{1}{\gamma^2}$$

$$\therefore d\mu = \frac{d\mu'}{\gamma^2 (1 + \beta \mu')^2}$$

and, since  $d\phi = d\phi'$

$$d\Omega = d\mu d\phi = \frac{d\mu' d\phi'}{\gamma^2 (1 + \beta \mu')^2} = \frac{d\Omega'}{\gamma^2 (1 + \beta \mu')^2}$$

Thus, the radiated energy per unit solid angle in the lab frame is  $dW = \gamma(1 + \beta \mu') dW' (\gamma^2 (1 + \beta \mu')^2)$

Thus, the radiated energy per unit solid angle in the lab frame is

$$\frac{dW}{d\Omega} = \frac{\gamma(1+\beta\mu')dW'}{d\Omega'} (\gamma^2(1+\beta\mu')^2)$$

$$= \gamma^3(1+\beta\mu')^3 \frac{dW'}{d\Omega'}$$

The ang. dist'n of the rest frame power is found by dividing  $\frac{dW'}{d\Omega'}$  by  $dt'$ . In the lab-frame the received

power requires including time-dilation; retardation effects

$$\text{i.e., } dt = \gamma(1-\beta\mu)dt' = \gamma\left(1-\beta\left[\frac{\mu'+\beta}{1+\beta\mu'}\right]\right)dt'$$

$$= \frac{dt'}{\gamma(1+\beta\mu')}$$

$$\therefore \frac{dP}{d\Omega} = \frac{dW}{d\Omega dt} = \frac{\gamma^3(1+\beta\mu')^3}{dt'} \frac{dW'}{d\Omega'} \gamma(1+\beta\mu')$$

$$\text{or } \frac{dP}{d\Omega} = \gamma^4(1+\beta\mu')^4 \frac{dP'}{d\Omega'}$$

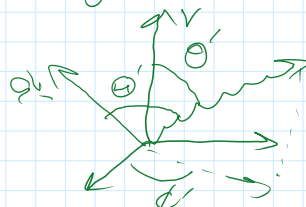
To get rid of  $\mu'$  on RHS, use  $\mu = \frac{\mu'+\beta}{1+\beta\mu'}$ ; flip signs and primes  $\mu' = \frac{\mu-\beta}{1-\beta\mu}$

ASA...

$$\frac{dP}{d\Omega} = \frac{1}{\gamma^4(1-\beta\mu)^4} \frac{dP'}{d\Omega'}$$

In the instantaneous rest frame of the particle, the ang. dist'n is given by a dipole (eg. Larmor)

$$\frac{dP'}{d\Omega'} = \frac{q^2 a'^2 \sin^2 \Theta'}{4\pi c^3}$$



Writing  $\vec{a}' = \vec{a}'_{\parallel} + \vec{a}'_{\perp}$  and using  $a'_{\parallel} = \gamma^3 a_{\parallel}$  ;  $a'_{\perp} = \gamma^2 a_{\perp}$

we get 
$$dP = \frac{q^2}{4\pi c^3} (\gamma^2 a_{\parallel}^2 + a_{\perp}^2) \sin^2 \Theta'$$

we get

$$\frac{dP}{d\Omega} = \frac{q^2}{4\pi c^3} \frac{(\gamma^2 a_{\parallel}^2 + a_{\perp}^2) \sin^2 \theta'}{(1 - \beta \mu)^4}$$